

SEVENTEEN RADIO TELESCOPES JOINTLY AND SIMULTANEOUSLY OBSERVE THREE QUASARS  
USING THE TECHNICAL PROCESS OF ELECTRONIC REAL-TIME VERY LONG BASELINE INTERFEROMETRY (e-VLBI)  
FOR 33 HOURS ON JANUARY 15-16 2009.

The participating telescopes included :

- 1) The 6m radio telescope at Observatorio Geodésico TIGO, Concepción, Chile
- 2) The 14m radio telescope at the Metsähovi Radio Observatory, Metsähovi, Kymälä, Finland
- 3) The 18.3m Westford Radio Telescope at MIT's Haystack Observatory, Westford, Massachusetts, USA
- 4) The 22m Mopra telescope at the Mopra Observatory, Coonabarabran, New South Wales, Australia
- 5) The six 22m antennas of the Australia Telescope Compact Array, Narrabri, New South Wales, Australia
- 6) The fourteen 25m antennas of the Westerbork Synthesis Radio Telescope, Dwingeloo, Netherlands
- 7) The 25m radio telescope at Onsala Space Observatory, Onsala, Sweden
- 8) The 25m radio telescope at the Urumqi Observatory, Nanshan, China
- 9) The 25m Sheshan radio telescope at the Shanghai Astronomical Observatory, Sheshan, China
- 10) The 26m Hobart radio telescope at Mt. Pleasant Observatory, Hobart, Tasmania, Australia
- 11) The 26m Mk2 telescope at Jodrell Bank Centre for Astrophysics, Lower Withington, UK
- 12) The 32m radio telescope at Cambridge, Jodrell Bank Centre for Astrophysics, Cambridge, UK
- 13) The 32m parabolic antenna of the Istituto Nazionale di Astrofisica, Medicina, Italy
- 14) The 32m precise parabolic antenna at the Torun Center for Astronomy, Piwnice, Poland
- 15) The 34m Kashima antenna of the Kashima Space Research center, Kashima, Japan
- 16) The 100m Radio Telescope Effelsberg, Max Planck Institute for Astronomy, Effelsberg, Germany
- 17) The 305m Arecibo radio telescope of the Arecibo observatory, Arecibo, Puerto Rico

Each telescope's readings were transmitted through global high-speed networks and combined by a dedicated supercomputer in the Netherlands. Some goals, according to internet sites about the project: 1) to "zoom in" on the most "energetic events" in the universe; 2) to utilize a REAL-TIME DISTRIBUTED ASTRONOMICAL INSTRUMENT OF INTERCONTINENTAL DIMENSIONS; 3) to continually observe a source for over 24 hours; 4) to obtain an image with over 1000x the resolution of one taken by the optical Hubble Space Telescope.

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